

IDENTIFICATION AND RECONSTRUCTION OF LOW-ENERGY ELECTRONS IN THE PROTODUNE-SP DETECTOR



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OUTLINE

- Introduction to low-energy electrons (Michel electrons)
- DUNE and ProtoDUNE
- Michel electron analysis
 - Event selection
 - Energy reconstruction
 - Missing energy studies
 - Energy resolution studies
 - Analysis current status
- Conclusion





LOW-ENERGY ELECTRONS STUDIES

- DUNE goal is to measure CP-violation via $v_{\rm e}$ appearance $(v_{\mu} \rightarrow v_{\rm e})$
 - To measure $v_{\rm e}$ appearance we need to select electron showers in LAr TPC, and reconstruct their energy
- Important to show that the detector can use the topological / calorimetric information provided by the TPC to identify a specific topology [Michel electrons]
- Reconstruct the low-energy electrons and produce their visible energy spectrum
- Ideal to study detector's response to electrons in the tens of MeV energy range
 - Useful for the search of low energy events e.g supernova

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 The goal is to deliver Michel analysis algorithm to use on DUNE-FD Day 1





INTRODUCTION TO DUNE

 The Deep Underground Neutrino Experiment (DUNE) is a future acceleratorbased multi-detector long-baseline neutrino oscillations experiment.



- Major DUNE goals:
 - Neutrino Oscillation Physics
 - Nucleon Decay
 - Supernova burst physics & astrophysics





INTRODUCTION TO DUNE

The Deep Underground Neutrino Experiment (DUNE) is a future acceleratorbased multi-detector long-baseline neutrino c^{--:-} experiment.



New neutrino beam at Fermilab,1300 km baseline

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- Multiple technologies for the Near Detector (ND) at Fermilab
- 70 kton Liquid Argon Time Projection Chamber (LArTPC) Far Detector (FD) at Sanford Underground Research Facility, South Dakota, 1.5 km underground
- Excavation started in 2017, begin taking data in late 2020s- early 2030s



PROTODUNE SP AND DP AT CERN

- Single-phase (SP) and dual-phase (DP) DUNE prototype LArTPCs at CERN
- 770 t LAr mass each
- Exposed to H2 (DP) and H4 (SP) test beams at CERN, momentum dependent beam composition contains *e*, K[±], μ[±], p, π[±]
- Also collected cosmic ray data
- ProtoDUNE-II is currently being commissioned







PROTODUNE-SP

- Active Volume: 6m (H) x 7m (L) x 2 x 3.6m (W)
- Central Cathode Plane Assembly (CPA) :
 - 3.6 m drift distance @180 kV
 - 500 V/cm field in drift volume
- Anode Plane Assembly (APA):
 - 3 APAs on each side
 - Each APA module: 6m high, 2.3m wide
 - Two induction planes and one collection plane



- Cold electronics: directly attached to the top of the APA (2560 wires/APA, 15360 total wires)
- Photon detectors (PDS): 3 designs integrated into APA frame bars
- Cryogenic Instrumentations outside of field cage: measure argon purity, temperature, liquid level and tag cosmic rays
- First papers from ProtoDUNE-SP are published e.g JINST 15 (2020) 12, P12004
- ProtoDUNE-SP Phase-I operated Sept. 2018 July 2020, Phase-II data taking under preparation

















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Collection plane











The third dimension is obtained by combining timing information (t₀) with drift velocity (v_d) → hence is called "Time projection chamber"

Light is fast compared to the electron drift!

Scintillation light is collected by PMTs that give absolute time for the interaction

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DUNE SENSITIVITY AND MEASURING DELTA-CP

- CP violation measured as a difference of neutrino and anti-neutrino oscillation probabilities. $P(v_{\mu} \rightarrow v_{e}) - P(\bar{v}_{\mu} \rightarrow \bar{v}_{e}) = -16s_{12}c_{12}s_{13}c_{13}^{2}s_{23}c_{2}\sin\delta\sin\left(\frac{\Delta m_{12}^{2}}{4E}L\right)\sin\left(\frac{\Delta m_{13}^{2}}{4E}L\right)\sin\left(\frac{\Delta m_{23}^{2}}{4E}L\right)$
- Calibrate Far and Near detector visible energy to 2% to control systematics at the level needed to support neutrino oscillation measurements
- Precision of energy reconstruction is required for
 - Early DUNE physics milestone is to reduce δ_{CP} uncertainty below 20° (at δ_{CP} = -90°)



ACCURATE RECONSTRUCTION OF VISIBLE ENERGY

- Perform oscillation analysis
 - Reduce systematics in lepton and neutrino interaction measurements
- Visible energy is obtained by **leptonic + hadronic** energy components
 - Goal is to achieve variation in visible energy to be <2% for DUNE sensitivity studies
- Working on leptonic part of the energy by:
 - Serving as Electromagnetic Shower working group convener at ProtoDUNE
 - Working on low-energy electron shower reconstruction
 - Developing calibration schemes for high energy electrons
- Hadronic energy component can be understood by neutrino interaction measurements





MICHEL ELECTRON ANALYSIS



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MICHEL ELECTRONS

- Michel electrons are electrons produced from the decay of muons (0-50 MeV)
- Common channels (in ProtoDUNE):
 - Muon decay: $\mu^+ \rightarrow e^+ \bar{\nu}_{\mu} \nu_e$ (80%)
 - Muon capture: $\mu \rightarrow e^{-} \nu_{\mu} \bar{\nu}_{e} n \gamma$ (20%)
- Unique features of this analysis:
 - data-driven determination of the recombination correction
 - evaluation of the missing Michel energy
 - comparison of the electron energy calibration based on muon derived calibration corrections with that based on the Michel electron true energy
 - characterization of the electron energy resolution







ANALYSIS FLOW









Position along the drift direction is only accurate for cathode crossers





























EVENT SELECTION Muon selection

Select candidate muon tracks with > 75 cm length

length > 75 cm



Sample purity = 27%



EVENT SELECTION Muon selection

- Select candidate muon tracks with > 75 cm length
- Minimum hit time > 200 ticks
 - -1 time tick = 500 nsec





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EVENT SELECTION Muon selection

- Select candidate muon tracks with > 75 cm length
- Minimum hit time > 200 ticks
- Maximum hit time < 5800 ticks</p>





Sample purity = 63%





Nearby hit count > 5 & < 40</p>





Sample purity = 82%





Nearby hit count > 5 & < 40</p>

Closest reco shower distance < 10 cm</p>





Sample purity = 87%





- Nearby hit count > 5 & < 40</p>
- Closest reco shower distance < 10 cm</p>
- Angle between muon and Michel < 130^o





Sample purity = 91%





- Nearby hit count > 5 & < 40</p>
- Closest reco shower distance < 10 cm</p>
- Angle between muon and Michel < 130^o
- Angle between Michel and collection plane wires
 > 10^o and < 170^o







MICHEL HITS OPTIMIZATION

- Form a cone that contains candidate Michel electron hits
- Look for the nearby charge deposition around the muon end position
- Optimize the parameters cone length (d) and angle (θ) to maximize Michel hit purity and hit completeness matrices

'Hit purity' = fraction of hits that belong to the true Michel 'hit completeness' = fraction of true Michel hits inside the







SYSTEMATIC UNCERTAINTY STUDIES

- Evaluated the impact of various systematics on the Michel energy spectrum
- The uncertainties are expressed with respect to the mean energy of the reconstructed Michel electron energy spectrum
- Some of the significant uncertainties are listed in the table

Uncertainty sources	Uncertainty estimates
Hit reconstruction efficiency systematic uncertainty	6.7%
Space charge effect systematic uncertainty	1.4%
Recombination factor systematic uncertainty	2.0%
Michel electron versus positron systematic uncertainty	1.7%
Total systematic uncertainties added in quadrature	7.3%



SPACE CHARGE SYSTEMATICS

- Reperformed the analysis from the MC sample generated with SCE OFF
- Compared Michel energy spectra with and without space charge
- The percentage difference in the mean of the distributions are taken to be the systematic uncertainty due to space charge





POSITRON VERSUS ELECTRON SYSTEMATICS

- Idea is to see whether Michel electrons behave differently than positron
- Plotted the true energy spectrums of positrons and electrons
- The difference in the mean of the distributions are taken to be the systematic uncertainty due to electron vs positron systematic





RECOMBINATION FACTOR SYSTEMATICS

 Evaluated the effective Michel recombination factor from MC sample using the modified box model

$$R = \frac{\ln\left(\frac{dE}{dx} \times \beta' / \rho E + \alpha\right)}{\frac{dE}{dx} \times \beta' / \rho E},$$

- This agrees to the recombination factor value used in the simulation within 2%
- The value is taken to be the systematic on recombination factor





HIT RECONSTRUCTION SYSTEMATICS

- Considered the number of hits closer to the muon end position
- Calculated the difference between data and MC distribution
- The value is taken to be the systematic on the hit reconstruction







CHARGE TO ENERGY CALIBRATION

Performed energy calibrations to extract Michel electron energy spectrum using modified box model*

$$E = \frac{C_{norm} * W_{ion}}{R * C_{calib}} * \sum_{i=1}^{N} \left[\varepsilon(X_i) * \varepsilon(Y_i, Z_i) * dQ_i \right]$$

- dQ_i is the reconstructed measured charge deposited on a hit
- W_{ion} is the work function for ionizing an argon atom
- C_{calib} represents the calibration constant that converts ADC to MeV
- C_{norm} normalizes the dQ/dx values to the dQ/dx at the anode plane
- ε(X_i), ε(Y_i,Z_i) are correction factors that are meant to remove the residual nonuniformities in the readout after correcting for all known effects (space charge, electron lifetime, etc).
- R = 0.6441 is recombination factor that accounts for recombination effects in the detector
- Calculation follows the same calibration procedure developed using the cosmic muons and applied to the beam data
 - "ProtoDUNE performance paper": <u>https://iopscience.iop.org/article/10.1088/1748-0221/15/12/P12004</u>





MICHEL ELECTRON RECONSTRUCTION

Muon-based method

- Simulation agrees very well with the data
 - Relative energy scales of data and Monte Carlo events agree to within 2%
- True versus reconstructed Michel energy plot shows a linear agreement



MICHEL ELECTRON RECONSTRUCTION

True-Michel based method

- Here, we match the shape of the reconstructed Michel charge directly with the true Michel energy smeared with a gaussian function
 - Optimize the gaussian uncertainty and charge to energy scale factor to find the best value of χ^2 /ndf between both distributions
- Simulation agrees very well with the scaled data
 - Relative energy scales of data and Monte Carlo events agree to within 3%
- True versus reconstructed Michel energy plot shows a linear agreement





MISSING ENERGY STUDIES Hit reconstruction threshold

- Investigated the source of missing energy due to hit reconstruction threshold
 - Applied at the signal processing level
- Since Michel electrons are already low in energy, losing a few hits significantly impact the final energy spectrum
- We lose 11% energy due to hit reconstruction threshold.







MISSING ENERGY STUDIES Hit incompleteness

- The other major source of missing energy are the hits outside the cone
- We introduce a lot of unwanted hits if we elongate the cone
- We lose 13% energy due to this effect

4 - ×10	³ DUNE:ProtoDUNE-SP Simulation		
4.5			
4			
3.5			
3			
ts	Outside cone		
U 2.5	energy per		
₩ 2 ₩	event		
1.5			
F			
E			
0.5			
0	2 4 6 8 10 12 14 16 18 20		
Outside Cone Michel energy [MeV]			

Energy	Percentage
Cone-only	76%
Off-cone	13%
Below hit reconstruction threshold	11%
Total	100%





FRACTIONAL ENERGY DIFFERENCE

Fractional energy difference is defined as

$$\Delta \varepsilon = \frac{(E_{true} - E_{reco})}{E_{true}}$$

 The Δε per event peaks at zero when we add both missing energy contributions to the simulation as expected







MICHEL ENERGY RESOLUTION



MICHEL ELECTRON PAPER STATUS

- We have drafted a technical note on this analysis
- The paper draft has been prepared that has completed its editorial board review
- Will soon be sent out to the entire DUNE collaboration for another round of review

Paper draft

Identification and reconstruction of low-energy electrons in the ProtoDUNE-SP detector (The DUNE Collaboration)

This article describes the selection and reconstruction of low energy electrons in the ProtoDUNE-SP detector. ProtoDUNE-SP is one of the prototypes for the Deep Underground Neutrino Experiment (DUNE) far detector, built and operated at CERN as a charged particle test beam experiment. The experiment collected data from August 2018 to July 2020. The analysis described here employs a fully automatic event selection and charged particle track reconstruction to obtain a sample of candidate cosmic muons, and to identify low-energy electrons around the end positions of selected candidate muons. Low-energy electron candidates are reconstructed using tools developed as a part of this analysis. Studies have been performed to verify the high purity (~95%) of selected low-energy electron event candidates and to calibrate the low-energy electron energy scale with both the cosmic-muon dE/dx method and the use of the theoretical low-energy electron energy spectrum. In addition, the effects of detector response to low-energy electrons including readout electronics threshold effects are quantified. Finally, the relation between the theoretical and reconstructed low-energy electron energy spectrum is derived and the energy resolution is characterized. Low-energy electron selection presented here accounts for about 75% of the total electron deposited energy. After the addition of missing energy the energy resolution improves from about 40% to 25% at 50 MeV.

I. INTRODUCTION

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Discoveries over the past half-century have positioned 57 18 neutrinos, one of the most abundant matter particles in 58 10 20 the universe, at the fundamental physics center stage. 59 Neutrinos are now being studied to answer open ques- 10 sembled and tested at the CERN Neutrino Platform [16], 21 tions about the nature of matter and the evolution of the 61 is a prototype of the first DUNE far detector module, 22 23 universe. In particular, the measurement of CP violation @ and it incorporates full-size components as designed for 24 in the lepton sector [1, 2] may help probe the possibility s1 that module. The ProtoDUNE-SP detector is a singlethat early-universe CP violation involving leptons might 4 phase LArTPC with an active volume of 7.2 × 6.1 × 25 have led to the present dominance of matter over anti- 55 7.0 m³. Major components of the ProtoDUNE-SP de-26 matter. Current long-baseline (LBL) neutrino oscillation 46 tector include the mechanical structure with a liquid ar-27 28 experiments are beginning to provide initial constraints or gon membrane cryostat, the time-projection chamber, the on the CP-violating phase δ_{CP} [3, 4], and definitive mea- ∞ photon detector system, electronics readout systems, and surements are planned to be performed by the dedicated 49 the computer data storage. ProtoDUNE-SP was operated 30 LBL accelerator neutrino experiments DUNE [5] and 70 from 2018 to 2020 and its collection of large samples 81 Hyper-Kamiokande [6]. The DUNE experiment [7] is 71 of high-quality beam data were used to demonstrate the 32 a next-generation LBL accelerator neutrino experiment, 72 effectiveness of the single-phase far detector design. Re $v_{\mu} \rightarrow v_{e}$ oscilla- v_{μ} sults on the performance of the ProtoDUNE-SP liquid ar-35 tions. The DUNE experiment will consist of a far de- 74 gon TPC in the test beam can be found in reference [17] tector [8] to be located about 1.5 km underground at π including noise and gain measurements, dE/dx calibra-37 the Sanford Underground Research Facility (SURF) in 76 tion for muons, protons, pions and electrons, drift elec-38 South Dakota, USA, at a distance of 1300 km from Fer- 77 tron lifetime measurements, and photon detector noise. milab, and a near detector [9] to be located at Fermi- 78 signal sensitivity and time resolution measurements. The 39 40 lab. The far detector will be a large modular liquid ar-79 measured values meet or exceed the specifications for the 41 gon time-projection chamber (LArTPC) with a total LAr 80 DUNE far detector. 42 mass of nearly 70 kt. With LAr technology it is possible 81 To achieve the planned DUNE physics program, it is as to reconstruct neutrino interactions with image-like nreason critically important to accurately reconstruct and mea-

54 low-energy electrons and positrons. Neutrino detectors at the far site of DUNE are planned to be built inside

four cryostats, each of which will contain 17.5 kt of liguid argon. The first detector to be constructed is planned to be a single-phase TPC [14].

The ProtoDUNE apparatus (ProtoDUNE-SP) [15] as-



SUMMARY

- Developed Michel electron selection, reconstruction, and energy calibration tools for ProtoDUNE-SP
 - Achieved 95% event selection purity
- Identified the major source of missing energy in the spectrum
 - Hit reconstruction threshold
 - Energy outside Michel reconstructed cone
- Presented Michel electron energy spectrum with two consistent methods
- Presented the Michel electron resolution
- A paper draft is under review within the collaboration
 Stay tuned!





THANK YOU FOR YOUR ATTENTION



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BACKUP SLIDES





PROTODUNE-SP—TERTIARY CERN SPS BEAM

- Tertiary CERN SPS Beam
 - Beam coming from Super Proton Synchrotron
- CERN SPS Beam:
 - Protons from 400 GeV CERN SPS
 - Incident on Beryllium Target
 - Secondary Beam is 80 GeV
 - Incident on secondary target
 - 0.3-7 GeV H4-VLE Beam





DUNE FD SP Prototype—Photon Detection System

Testing new technologies!

ProtoDUNE-SP consists of:

- Tertiary CERN SPS Beam
- Single-Phase LArTPC
- Photon Detection System (PDS)

127 nm LAr scintillation light

-430 nm shifted light (in surface)

• Cosmic-Ray Tagger (CRT)

"Dip-Coated" (DC) Type Collector × 29

"Double-Shifted" (DS) Type Collector \times 29



Photon Detection System

- Photon Detector Technology Demonstration
 - 60 Photon Detectors
 - 3 Photon Collector Technologies
 - 29 Dip Coated Light Guide Detectors
 - 29 Double Shift Light Guide Detectors
 - 2 S-ARAPUCAs
 - 2 Photon Sensors Manufactures
 - SensL & Hamamatsu



LOW-ENERGY ELECTRON STUDIES AT DUNE FAR DETECTOR

- Developing selection and reconstruction framework to isolate muons and its Michel electrons to calibrate electron energy scale at Far Detector (10 kt module).
 - Achieved 80% event purity
 - True Michel efficiency is 31%
 - My plan is to deliver Michel analysis algorithm for use in Far Detector on day one



PASSING RATE AS A FUNCTION OF TRUE MICHEL ENERGY

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